

Lecture 10

Last time wrote \mathcal{L} :

$$\mathcal{L} = -\frac{1}{4} Z_3 F_{\mu\nu}^2 - i Z_2 \bar{\psi} \not{\partial} \psi - Z_m m \bar{\psi} \psi \\ - i Z_1 \bar{\psi} \not{A} \psi$$

When $Z_i = 1 + O(e^2)$ & chosen to cancel off divergences arising from loops

Furthermore physical observables computed using this \mathcal{L} are μ indep.

(artificial scale arising in dim reg)

truly rewrite as

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu}^2 - i \bar{\psi}_0 \not{\partial} \psi_0 - m_0 \bar{\psi}_0 \psi_0 \\ - i \bar{\psi}_0 \not{A}_0 \psi_0 \quad \begin{array}{l} \uparrow \quad \uparrow \\ \text{bare parameters /} \\ \text{fields} \end{array}$$

m_0, ψ_0 , etc indep of μ , likely so but

yield finite observables!

Trivially

$$A_0 = \sqrt{Z_3} A$$

$$\phi_0 = \sqrt{Z_2} \phi$$

$$m_0 = z_m / z_2^m$$

$$e_0 = \left(\frac{z_1}{z_2 \sqrt{z_3}} \right) e \mu^{\epsilon/2}, \quad \epsilon = 4-d.$$

notice: A, ϕ, m finite + renormalized quantities (cut-off indep) but in general they do depend on μ .

but m_0, A_0, ϕ_0 are μ indep (but likely ∞ !)

$$z_1 = 1 - e^2 / 8\pi^2 \frac{1}{\epsilon}$$

$$z_2 = 1 - e^2 / 8\pi^2 \frac{1}{\epsilon}$$

$$z_3 = 1 - e^2 / 6\pi^2 \frac{1}{\epsilon}$$

$$z_m = 1 - e^2 / 2\pi^2 \frac{1}{\epsilon}$$

$z_1 = z_2$ again

ignoring

finite

parts

e_0 MS scheme

Taking logs: \Rightarrow

$$\ln e_1 = \ln \left(\frac{z_1}{z_2 \sqrt{z_3}} \right) + \ln e + \frac{\epsilon}{2} \ln \mu.$$

$$= \frac{e^2}{12\pi^2 \epsilon} + \ln e + \frac{\epsilon}{2} \ln \mu.$$

$$\frac{\partial \ln e_1}{\partial \ln \mu} \Rightarrow \Rightarrow$$

$$0 = \left(1 + \frac{1}{\epsilon} \frac{e^2}{6\pi^2} \right) \frac{\partial e}{\partial \ln \mu} + \frac{\epsilon e}{2}$$

imply

$$\frac{\partial e}{\partial \ln \mu} = -\frac{\epsilon e}{2} + \beta(e)$$

\uparrow Beta function

$$\left(-\frac{\epsilon e}{2} + \beta \right) \left(1 + \frac{1}{\epsilon} \frac{e^2}{6\pi^2} + \dots \right) = 0$$

terms of order $\epsilon^0 \rightarrow$

$$\beta(e) - \frac{e^3}{12\pi^2} = 0$$

tells you how to ~~adjust~~ renormalized charge must vary with parameter μ to keep physical observables μ independent.

Order: ^{about} what the finite parts in Z_i 's?

eg β function repeat with ~~$f(\mu)$~~

$$\text{eg } Z_i = 1 - e^2 / 8\pi^2 \frac{1}{\epsilon} + f(\mu)$$

$$(1 + \frac{1}{\epsilon} e^2 / 6\pi^2 + f(\mu)) (-\epsilon e / 2 + \beta)$$

$f(\mu)$ does not contribute at $O(\epsilon^0)$

\hookrightarrow ~~β~~ , γ independent of finite parts
in $\boxed{d=4}$ ($\epsilon \rightarrow 0$)

Explains why $\overline{\text{MS}}$ scheme so attractive
Easiest to just get divergent pieces from
Feynman integrals. All one needs for β , γ etc.

(disadvantages: p deriv propagator not linear
at physical fermion mass ...)

β function ≥ 0

charge increases with energy scale μ .

(Landau pole again)

Can apply same trick to other

renormalized / bare parameters eg mass

$$\ln m_0 = \ln(Z_m/Z_2) + \ln m$$

$$= -\frac{e^2}{8\pi^2} \frac{1}{\epsilon} + \frac{e^2}{8\pi^2} \frac{1}{\epsilon} + \ln m$$

$$\frac{\partial \ln m_0}{\partial \ln \mu} = 0 \Rightarrow -\frac{3}{8\pi^2} \frac{1}{\epsilon} \frac{2e\partial e}{\partial \ln \mu} + \frac{1}{m} \frac{\partial m}{\partial \ln \mu} = 0$$

$$\text{now } \frac{\partial e}{\partial \ln \mu} = -\epsilon e/2 + \beta$$

define
mas anomalous dimension $\gamma_m = \frac{1}{m} \frac{\partial m}{\partial \ln \mu}$

$$\Rightarrow \gamma_m = \frac{3}{4\pi^2} \frac{1}{\epsilon} e (\beta - \epsilon e/2)$$

at $O(\epsilon^0) \Rightarrow$

$$\gamma_m = \frac{-3}{8\pi^2} e^2 + O(e^4)$$

Similarly can define anomalous dimensions of fields ψ, A

$$\frac{1}{2} \partial / \partial \ln \mu \quad Z_2 = \gamma_\psi(e)$$

$$\frac{1}{2} \partial / \partial \ln \mu \quad Z_3 = \gamma_A(e)$$

Furthermore bare Green functions

$$\Gamma_0^{(n)} = \langle \psi_0 \dots \bar{\psi}_0 \dots A_0 \rangle$$

are also indep of μ

use this fact to derive differential eqs

renormalized Green functions

Callan-Symanzik eqs (RGE's)

$$\text{eg } \langle \bar{\psi}_0 \psi_0 \rangle = \Delta_0(k^2)$$

$$\frac{\partial}{\partial \ln \mu} \ln \Delta_0(k^2) = 0$$

$$\text{but } \Delta_0(k^2) = Z_2 \Delta(k^2)$$

$$\therefore 0 = \frac{\partial \ln Z_2}{\partial \ln \mu} + \frac{1}{\Delta(k^2)} \left(\frac{\partial}{\partial \ln \mu} + \frac{\partial \ell}{\partial \ln \mu} \frac{\partial}{\partial \ell} + \frac{\partial M}{\partial \ln \mu} \frac{\partial}{\partial M} \right) \Delta(k^2)$$

since Δ depends on e, m & μ .

$$\text{or } \left(\frac{\partial}{\partial \ln \mu} + \beta(e) \frac{\partial}{\partial e} + \gamma_m m \frac{\partial}{\partial m} + 2\gamma_\psi \right) \Delta(k^2) = 0$$

notice : if $m \rightarrow 0$ & $\beta = 0$ (LFT)

$$\left(\frac{\partial}{\partial \ln \mu} + 2\gamma_\psi \right) \Delta = 0$$

~~$$\Delta \sim \mu^{-2\gamma_\psi} = (k^2)^{-\gamma}$$~~

$$\Delta \sim \mu^{-2\gamma_\psi} = (\mu^2)^{-\gamma}$$

or putting in dimensions

$$\Delta \sim (\mu/k)^{-2\gamma_\psi}$$

but we know

$$\Delta(k^2) \sim \frac{1}{k^2}$$

(insert in diff eq for wrt μ)

$$\Delta \sim \frac{1}{k^2} \left(\frac{\mu^2}{k^2} \right)^{-\gamma}$$

\bar{c}
 $2 \Rightarrow 2 - 2\gamma$
 under k scaling

Why the term anomalous dimensions?

Consider scalar field theory

$$S = \int d^4x (\phi \square \phi + m^2 \phi^2 + \lambda \phi^4)$$

Invariant under dilation / scale symmetry:

$$x \rightarrow \lambda x \quad \partial \rightarrow \lambda^{-1} \partial \quad m \rightarrow \lambda^{-1} m \quad g \rightarrow g \quad \phi \rightarrow \lambda^{-\alpha} \phi$$

Same argument error.

$$G^n = \langle \phi_1 \dots \phi_n \rangle \rightarrow \lambda^{-n\alpha} G^n \quad \text{classically}$$

however on the side μ appears in renormalized G^n 's.

$$G^n(x, g, m, \mu) \sim m^a g^b x_1^{c_1} \dots x_n^{c_n} \mu^\delta$$

classically $a - c_1 - \dots - c_n = n$ classical scaling dimension of G^n .

$$qm: a + \delta - c_1 - \dots - c_n = n$$

$$\Rightarrow a - c_1 - \dots - c_n = \underline{n - \delta}$$

thus under dilation $G^n \rightarrow \lambda^{n-\delta} G^n$

since μ does not ~~scale~~ transform under dilation?

thus classical scaling dim shifts $n \rightarrow n - \delta$
 $\Rightarrow \partial G^n / \partial \ln \mu = \delta$ as above

Further comments on renormalization

RG

- * QED is an example of a renormalizable QFT
- * This is just the statement that a finite # of counterterms are needed to remove all d/s to all orders of p. theory
- * Standard arguments tell you that this property is tied to the fact that the L contains no interactions with negative mass dimension couplings
- * Actually \Rightarrow 4d QED (\Rightarrow is non-abelian (causing) pretty much saturate the class of renormalizable QFT (at least perturbatively)

* Key idea: quantum effects can strongly affect values of parameters in classical Lagrangian

Indeed the classical parameters are in principle unobservable & the best we can hope to do is to reparameterize theory in terms of renormalized parameters whose values are ultimately taken from expt.

* Problems all stem from short distance scales (UV) & an equivalent way to think of the procedure of renormalization is that it is a systematic way of ~~removing~~ renumbering parameters of theory in ~~way~~ such a way as to render low energy observables insensitive to UV scales

* A physicist at one energy scale should not be sensitive to physics at some much higher scale ~~####~~

Price you pay

* Removing this structure to U-V scales/cutoff, leaves a remaining imprint: the coupling constants are not constant!

They change with physical scale (μ)

$$\ln \mu^2 \rightarrow \ln \mu^2/k^2$$

* This is seen most graphically in

the running of the interaction strength with momentum scale $e(\mu)$ or $e(k)$

— β function etc

anomalous scaling of correlation functions via γ .

Important part

* Many of these comments also apply to so-called non-renormalizable theories

* In principle such theories require a ∞ # counter terms if one wants to work to all orders in the coupling.

However to low order the non-renormalizable interactions only contribute terms $O\left(\frac{E}{\Lambda}\right)^n$ where Λ is some scale needed to write down the non-renormalizable operator in the classical theory

* Perturbation theory breaks down when $E \rightarrow \Lambda$ but if $E \ll \Lambda$ the theory can yield precise predictions for physical observables including loop effects.

"effective field theories"

Wilson's condensed matter

- * Indeed this is precisely what happens in C.M.P. In that case one is interested in phenomena at distance scales much greater than the lattice spacing ($\frac{1}{\Lambda}$)
- * In this case details of \mathcal{L} at short distances including all sorts of non-renormalizable ops are irrelevant to the behavior of system at large scales (hence term irrelevant ops.)
- * Related to concept of Universality: different microscopic \mathcal{L} all lead to same long distance physics (controlled by the renormalizable or relevant ops)
- * Wilson made all this clear and developed an alternative to the continuum RG which is in many ways more intuitive & plausible ...

Wilson's key idea

- ① Think of you theory as a lattice with cut-off $\Lambda = 1/a$
- ② Long distance physics should be insensitive to Λ
- ③ You should be able to change the "bare" parameters of theory to compensate for changes in $\Lambda \rightarrow$ leave long distance physics invariant
- ④ In practice this long distance universality only occurs if one is on a critical surface in space of bare parameters. (corresponds to tuning all relevant parameters eg masses to ~~the~~ special values (e.g. zero!))
- ⑤ On such a critical surface there may occur fixed pts ^{where} ~~where~~ physics is scale invariant
 $\beta = 0$ - CFT's

* Critical exponents in CIP are related to γ 's in QFT & characterize the ^{RG} flows close to one of these fixed pts

* Wilson & others developed series of techniques

1. ϵ Epsilon-expansion: keep ~~$d=4$~~ $d=4$ finite & expand $d \rightarrow d=3$ eg Wilson-Fisher

2. exact RG equations express Λ indep of physical observables (later?)

3. Numerical / Monte Carlo implementations of RG in real space

← Powerful techniques which have taught us a lot about strongly interacting QFTs